

Monte-Carlo Radiation Transport

Refs: arXiv:2103.16588, arXiv:2209.02538

April 17, 2025



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Outline

1 Monte-Carlo Radiation Transport

- Distribution Function
- Propagation of Packets
- Collision Terms
- Stress-energy Tensor and Moments
- Neutrino-matter Coupling

2 AMReX Particle Interface

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Distribution Function

When evolving the equations of radiation transport in general relativistic simulations, we aim to determine the distribution function of particles

$$f(t, x^i, p^\mu) = \sum_k \delta^3(x^i - x_k^i(t)) \delta^3(p_i - p_i^k(t)). \quad (1)$$

The distribution function follows Boltzmann's equation of radiation transport

$$p^\alpha \left[\frac{\partial f}{\partial x^\alpha} - \Gamma^\beta_{\alpha\gamma} p^\gamma \frac{\partial f}{\partial x^\beta} \right] = \left[\frac{df}{d\tau} \right]_{\text{collisions}}. \quad (2)$$

- Evolve a six-dimensional function in time

Monte Carlo Methods

Discretize the distribution function of neutrinos using Monte-Carlo packets, each representing a large number of neutrinos, i.e.

$$f(t, x^i, p^\mu) = \sum_{k=0}^{n_p} N_k \delta^3(x^i - x_k^i(t)) \delta^3(p_i - p_i^k(t)) \quad (3)$$

- Each packet represent N_k particles and has a single position and momentum

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Propagation of Packets

Neglecting the finite mass of neutrinos, we evolve packets along null geodesics (Hughes et al. 1994),

$$\frac{dx^i}{dt} = \gamma^{ij} \frac{p_j}{p^t} - \beta^i, \quad (4)$$

$$\frac{dp_i}{dt} = -\alpha p^t \partial_i \alpha + p_j \partial_i \beta^j - \frac{1}{2} p_j p_k \partial_i \gamma^{jk}, \quad (5)$$

for a null vector, $p^t = \frac{\sqrt{\gamma^{ij} p_i p_j}}{\alpha}$.

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1 Monte-Carlo Radiation Transport

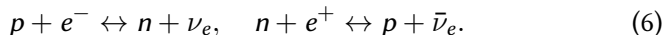
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Tabulated Reaction Rate

Use the **NuLib** library (O'Connor & Ott 2010) to generate tabulated values of the **emissivity** η , **absorption opacity** κ_a , and **elastic scattering opacity** κ_s experienced by Monte Carlo packets.

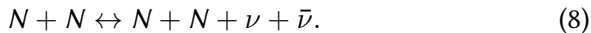
- Charged current reactions



- Pair production/annihilation



- Nucleon-nucleon bremsstrahlung



The output of the NuLib library, is a 4D table for η , κ_a and κ_s as a function of **fluid density** ρ , **fluid temperature** T , **fluid electron fraction** Y_e , and **neutrino energy** ν .

Emission

The total energy of the emitted neutrinos of that species and in that energy bin is

$$E_{\text{tot}} = \sqrt{-g} V \Delta t \eta, \quad (9)$$

If the desired energy of neutrino packets within this cell is E_{target} and the central value of the neutrino energies in our energy bin is ν ,

- we emit, on average, $E_{\text{tot}}/E_{\text{target}}$ packets, each representing E_{target}/ν neutrinos,
- the location of the packet is randomly drawn from a homogeneous distribution *in the coordinates of the simulation*,
- the 4-momentum of the neutrinos is drawn from an isotropic distribution in the fluid frame,

$$p_{\text{fl}}^{\hat{\mu}} = \nu(1, \sin \theta \cos \phi, \sin \theta \sin \phi, \cos \theta). \quad (10)$$

Absorption and Elastic Scattering

Before propagating a neutrino, we draw random numbers r_a, r_s from a homogeneous distribution in $[\epsilon, 1)$ (with $\epsilon = 10^{-70}$). The time to the next absorption/scattering event is then

$$\Delta t_{a,s} = -\frac{\kappa_{a,s} p^t}{\nu} \ln r_{a,s}, \quad (11)$$

Consider the smallest time interval between $\Delta t_{a,s}$ and the desired time step

- If Δt_a is the smallest, the packet is absorbed.
- If Δt_s is the smallest, the packet is scattered.
 - randomly draw a new 4-momentum with the same fluid-frame energy as the original packet
 - a direction of propagation drawn from an isotropic distribution

This simple process works well as long as $\kappa_{a,s} \Delta t \lesssim 1$, i.e., when individual grid cells are optically thin or semitransparent.

Pair Annihilation in Low-density Regions

- The absorption opacity for pair annihilation

$$\kappa_p = \frac{C_{\text{pair}} c C_F^2}{3\pi} \frac{p^\alpha p^\beta}{\nu} \bar{T}_{\alpha\beta}. \quad (12)$$

$$\bar{\kappa}_p = \frac{C_{\text{pair}} c C_F^2}{3\pi} \frac{\bar{p}^\alpha \bar{p}^\beta}{\bar{\nu}} T_{\alpha\beta}. \quad (13)$$

- This estimate of κ_p is only valid in low-density regions, we choose to suppress κ_p in dense regions,

$$\kappa_p \rightarrow \kappa_p e^{-\rho/\rho_{\text{crit}}}. \quad (14)$$

Pair Annihilation in Low-density Regions

- We take κ_p into account by correcting the number of neutrinos N represented by any given packet at the end of a time step

$$\frac{dN}{dt} = -\kappa_p N \rightarrow N(t) = N_0 e^{-\kappa_p(t-t_0)}. \quad (15)$$

After a time step Δt , we thus have $N(t_0 + \Delta t) = N(t_0)e^{-\kappa_p \Delta t}$.

- Pair annihilation is treated differently from other absorption processes: it never destroys Monte Carlo packets, but **only changes the number of neutrinos in each packet**.

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Stress-energy Tensor and Moments

The general relativistic stress-energy tensor is,

$$T_{\mu\nu}(t, x^i) = \sum_{k=1}^{n_p} N_k \frac{p^k_{\mu} p^k_{\nu}}{\sqrt{-g} p_k^t} \delta^3(x^i - x_k^i(t)). \quad (16)$$

An observer with 4-velocity u^{μ} measures the corresponding energy density

$$J = T_{\mu\nu} u^{\mu} u^{\nu} = \sum_{k=1}^{n_p} N_k \frac{\nu_k^2}{\sqrt{-g} p_k^t} \delta^3(x^i - x_k^i(t)), \quad (17)$$

with $\nu_k = -p^k_{\mu} u^{\mu}$ the energy of neutrinos in packet k as measured by our observer.

Stress-energy Tensor and Moments

The average energy density within a region of coordinate volume V is then

$$J = \sum_{k \in V} N_k \frac{\nu_k^2}{\sqrt{-g} V p_k^t}. \quad (18)$$

Similarly, the average linear momentum $H_\alpha = -T^{\mu\nu} u_\mu (g_{\nu\alpha} + u_\nu u_\alpha)$ measured by an observer with u^μ is

$$H_\alpha = \sum_{k \in V} N_k \frac{\nu_k (p_\alpha^k - \nu_k u_\alpha)}{\sqrt{-g} V p_k^t}. \quad (19)$$

Stress-energy Tensor and Moments

To couple particles with the fluid, we will also need to compute terms of the form

$$\int dt \kappa J = \sum_{k=1}^{n_p} \kappa_k N_k \frac{\nu_k^2}{\sqrt{-g} p_k^t} \Delta t_k \delta^3(x^i - x_k^i(t)), \quad (20)$$

with κ an opacity that depends on position and momentum of a particle. The average of this integral within a grid cell of volume V ,

$$\int dt \kappa J = \sum_{k \in V} \kappa_k N_k \frac{\nu_k^2}{\sqrt{-g} V p_k^t} \Delta t_k. \quad (21)$$

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Source Terms for Fluid Evolution

The source terms appearing in the evolution of these variables due to neutrino-matter interactions are

$$\partial_t \tilde{\tau} = \dots + \alpha \sqrt{\gamma} S^\alpha n_\alpha, \quad (22)$$

$$\partial_t \tilde{S}_i = \dots - \alpha \sqrt{\gamma} S^\alpha \gamma_{\alpha i}, \quad (23)$$

$$\partial_t (\rho_* Y_e) = \dots - \sum s_i m_p \alpha \sqrt{\gamma} \frac{\eta - \kappa_a J}{\nu}. \quad (24)$$

The source term is,

$$S^\alpha = \sum \eta' u^\alpha - \sum (\kappa'_a J u^\alpha + (\kappa'_a + \kappa'_s) H^\alpha). \quad (25)$$

Calculating the changes in the evolved fluid variables over one time step Δt thus requires calculations of

$$\int \kappa'_a J dt; \quad \int (\kappa'_a + \kappa'_s) H^\alpha dt; \quad \int \kappa'_a J dt / \nu, \quad (26)$$

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Main Feature

- Particle Packets: $(x^i, p^\mu, N_k, \text{species}, \dots)$,
- Propagation of Packets,
- Emission, Absorption, Scattering, ...,
- Interaction with Fluid.

Particle and ParticleContainer

```
Particle<2, 2> p;  
  
p.pos(0) = 1.0;  
p.pos(1) = 2.0;  
p.pos(2) = 3.0;  
p.id() = 1;  
p.cpu() = 0;  
  
// p.rdata(0) is the first extra real component, not the  
// first real component overall  
p.rdata(0) = 5.0;  
p.rdata(1) = 5.0;  
  
// and likewise for p.idata(0);  
p.idata(0) = 17;  
p.idata(1) = -64;
```

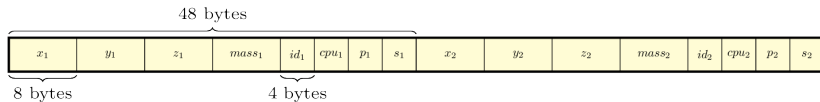
Particle and ParticleContainer

```
using MyParticleContainer = ParticleContainer<3, 2, 4, 4>;  
MyParticleContainer mypc;
```

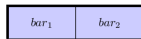
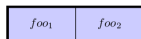
- For the first two template parameters, particles added to the container are stored in the Array-of-Structs style.
- There are two more optional template parameters that allows the user to specify additional particle variables that will be stored in Struct-of-Array form.

Arrays-of-Structs and Structs-of-Arrays

Array-of-Structs



Struct-of-Arrays



8 bytes



4 bytes

Constructing ParticleContainers

```
ParticleContainer (const Geometry           & geom,  
                  const DistributionMapping & dmap,  
                  const BoxArray          & ba);
```

```
ParticleContainer (const Vector<Geometry>           & geom,  
                  const Vector<DistributionMapping> & dmap,  
                  const Vector<BoxArray>            & ba,  
                  const Vector<int>                 & rr);
```

Redistribute

After calling `Redistribute()`

- all the particles will be moved to their proper places in the container,
- all invalid particle (particles with id set to -1) will be removed.

Initializing Particle Data

```
for (MFIter mfi = MakeMFIter(lev); mfi.isValid(); ++mfi) {  
  
    // ``particles'' starts off empty  
    auto& particles = GetParticles(lev)[std::make_pair(mfi.index(),  
                                                       mfi.LocalTileIndex())];  
  
    ParticleType p;  
    p.id()      = ParticleType::NextID();  
    p.cpu()     = ParallelDescriptor::MyProc();  
    p.pos(0) = ...  
    etc...  
  
    // AoS real data  
    p.rdata(0) = ...  
    p.rdata(1) = ...  
  
    // AoS int data  
    p.idata(0) = ...  
    p.idata(1) = ...  
}
```

Initializing Particle Data

```
// Particle real attributes (SoA)
std::array<double, 2> real_attribs;
real_attribs[0] = ...
real_attribs[1] = ...

// Particle int attributes (SoA)
std::array<int, 2> int_attribs;
int_attribs[0] = ...
int_attribs[1] = ...

particles.push_back(p);
particles.push_back_real(real_attribs);
particles.push_back_int(int_attribs);

// ... add more particles if desired ...
}
```

Iterating over Particles

- Iterate over all the AoS data

```
using MyParConstIter = MyParticleContainer::ParConstIterType;
for (MyParConstIter pti(pc, lev); pti.isValid(); ++pti) {
    const auto& particles = pti.GetArrayOfStructs();
    for (const auto& p : particles) {
        // do stuff with p...
    }
}
```

- Access the SoA data

```
using MyParIter = MyParticleContainer::ParIterType;
for (MyParIter pti(pc, lev); pti.isValid(); ++pti) {
    auto& particle_attributes = pti.GetStructOfArrays();
    RealVector& real_comp0 = particle_attributes.GetRealData(0);
    IntVector& int_comp1 = particle_attributes.GetIntData(1);
    for (int i = 0; i < pti.numParticles; ++i) {
        // do stuff with your SoA data...
    }
}
```

Interacting with Mesh Data

```
Ex.FillBoundary(gm.periodicity());
Ey.FillBoundary(gm.periodicity());
Ez.FillBoundary(gm.periodicity());
for (MyParIter pti(MyPC, lev); pti.isValid(); ++pti) {
    const Box& box = pti.validbox();

    const auto& particles = pti.GetArrayOfStructs();
    int nstride = particles.dataShape().first;
    const long np = pti.numParticles();

    const FArrayBox& exfab = Ex[pti];
    const FArrayBox& eyfab = Ey[pti];
    const FArrayBox& ezfab = Ez[pti];

    interpolate_cic(particles.data(), nstride, np,
                   exfab.dataPtr(), eyfab.dataPtr(), ezfab.dataPtr(),
                   box.loVect(), box.hiVect(), plo, dx, &ng);
}
```

Interacting with Mesh Data

```
rho.setVal(0.0, ng);  
for (MyParIter pti(*this, lev); pti.isValid(); ++pti) {  
    const Box& box = pti.validbox();  
  
    const auto& particles = pti.GetArrayOfStructs();  
    int nstride = particles.dataShape().first;  
    const long np = pti.numParticles();  
  
    FArrayBox& rhofab = (*rho[lev])[pti];  
  
    deposit_cic(particles.data(), nstride, np, rhofab.dataPtr(),  
                box.loVect(), box.hiVect(), plo, dx);  
}  
  
rho.SumBoundary(gm.periodicity());
```